

# Impact of fermionic singlets on lepton universality tests

Cédric Weiland

Laboratoire de Physique Théorique d'Orsay, Université Paris-Sud 11, France

Fermilab  
Batavia, June 13th, 2013



# Neutrino oscillations

- Oscillation probability is non-zero only if  $\Delta m_{kj}^2 = m_k^2 - m_j^2$  and mixing angles are different from zero

$$P_{\nu_\alpha \rightarrow \nu_\beta} = \delta_{\alpha\beta} - 4 \sum_{k>j} \Re[U_{\alpha k}^* U_{\beta k} U_{\alpha j} U_{\beta j}^*] \sin^2 \left( \frac{\Delta m_{kj}^2 L}{4E} \right) + 2 \sum_{k>j} \Im[U_{\alpha k}^* U_{\beta k} U_{\alpha j} U_{\beta j}^*] \sin \left( \frac{\Delta m_{kj}^2 L}{2E} \right)$$

- Neutrino oscillations:

- solar  $\nu_e \rightarrow \nu_{\text{others}}$ :  $\theta_{12} \simeq 33^\circ$ ,  $\Delta m_{12}^2 \simeq 7.5 \times 10^{-5} \text{eV}^2$  (best fit)
- atmospheric  $\overset{(-)}{\nu_\mu} \rightarrow \overset{(-)}{\nu_\tau}$ :  $\theta_{23} \simeq 40^\circ$  or  $50^\circ$ ,  $|\Delta m_{23}^2| \simeq 2.4 \times 10^{-3} \text{eV}^2$  (best fit)
- reactor  $\bar{\nu}_e \rightarrow \bar{\nu}_{\text{others}}$ :  $\theta_{13} \simeq 8.7^\circ$  (best fit)
- accelerator  $\nu_\mu \rightarrow \nu_{\text{others}}$



# Neutrino oscillations

- Oscillations  $\Rightarrow$  **Non-diagonal** charged currents

$$\mathcal{L}_{\text{int}} = -\frac{g}{\sqrt{2}} \mathbf{U}_\nu^{ji} \bar{\ell}_j \gamma^\mu P_L \nu_i W_\mu^- + \text{h.c.}$$

- 3 mass eigenstates  $\nu_i = \nu_1, \nu_2, \nu_3$  different from the interaction eigenstates  $\nu_\alpha = \nu_e, \nu_\mu, \nu_\tau$

$$\nu_\alpha = \mathbf{U}_\nu^{\alpha i} \nu_i$$

$\Rightarrow \mathbf{U}_\nu$  is a  $3 \times 3$  unitary matrix,  $U_\nu = U_{\text{PMNS}}$

- $U_{\text{PMNS}}$ : 3 mixing angles + 1 CP violating phase (+ 2 Majorana phases)
- Oscillations give no information on:
  - the absolute mass scale
  - the Dirac or Majorana nature of neutrinos



# Neutrino masses and nature

- Absolute masses

- Tritium  $\beta$  decays:  $m_{\nu_e} < 2.05$  eV [Kraus et al., 2005, Aseev et al., 2011]
- Cosmology: CMB  $\Sigma m_{\nu_i} < 0.98$  eV Planck: [Ade et al., 2013]  
CMB+BAO+ $H_0$ +flat Universe  $\Sigma m_{\nu_i} < 0.23$  eV

- Hierarchy:  $m_1 < m_2 < m_3$  or  $m_3 < m_1 < m_2$  ?

- Matter effect in the Sun:  $\nu_1$  is mostly  $\nu_e$
- $\text{sign}(\Delta m_{23}^2)$  ?  $\rightarrow$  matter effect in long baseline oscillations (T2K, NO $\nu$ A, future projects)

- Neutrino nature (Dirac or Majorana):

- Neutrinoless double  $\beta$  decay:  
 $m_{2\beta} < 0.12 - 0.25$  eV KamLAND-ZEN: [Gando et al., 2013]



# Massive neutrinos and New Physics

- Standard Model  $L = \begin{pmatrix} \nu_L \\ \ell_L \end{pmatrix}, H = \begin{pmatrix} H^+ \\ H^0 \end{pmatrix}$ 
  - No right-handed neutrino  $\nu_R \rightarrow$  No Dirac mass term

$$\mathcal{L}_{\text{mass}} = -Y_\nu \bar{L} \tilde{H} \nu_R + \text{h.c.}$$

- No Higgs triplet  $T \rightarrow$  No Majorana mass term

$$\mathcal{L}_{\text{mass}} = -\frac{1}{2} m \bar{L}^C T L + \text{h.c.}$$

- Necessary to go beyond the Standard Model for  $\nu$  mass
  - Radiative models
  - Extra-dimensions
  - R-parity violation in supersymmetry
  - Seesaw mechanisms  $\rightarrow \nu$  mass at tree-level  
+ BAU through leptogenesis



# Dirac neutrinos ?

- Add **gauge singlet**, right-handed neutrinos  $\nu_R$

$$\Rightarrow \nu = \nu_L + \nu_R$$

$$\mathcal{L}_{\text{mass}}^{\text{leptons}} = -Y_\ell \bar{L} H \ell_R - Y_\nu \bar{L} \tilde{H} \nu_R + \text{h.c.}$$

$\Rightarrow$  After electroweak symmetry breaking  $\langle H \rangle = \begin{pmatrix} 0 \\ v \end{pmatrix}$

$$\mathcal{L}_{\text{mass}}^{\text{leptons}} = -v Y_\ell \bar{\ell}_L \ell_R - v Y_\nu \bar{\nu}_L \nu_R + \text{h.c.} = -m_\ell \bar{\ell}_L \ell_R - m_D \bar{\nu}_L \nu_R + \text{h.c.}$$

$\Rightarrow$  **3** light neutrinos:  $m_\nu \lesssim 1\text{eV} \Rightarrow Y^\nu \lesssim 10^{-11}$

- Increase the hierarchy between Yukawa couplings



# Majorana neutrinos ?

- Add **gauge singlet**, right-handed neutrinos  $\nu_R$

$$\mathcal{L}_{\text{mass}}^{\text{leptons}} = -Y_\ell \bar{L} H \ell_R - Y_\nu \bar{L} \tilde{H} \nu_R - \frac{1}{2} M_R \overline{\nu_R^C} \nu_R + \text{h.c.}$$

$\Rightarrow$  After electroweak symmetry breaking  $\langle H \rangle = \begin{pmatrix} 0 \\ v \end{pmatrix}$

$$\mathcal{L}_{\text{mass}}^{\text{leptons}} = -m_\ell \bar{\ell}_L \ell_R - m_D \bar{\nu}_L \nu_R - \frac{1}{2} M_R \overline{\nu_R^C} \nu_R + \text{h.c.}$$

$\Rightarrow$  **6** mass eigenstates:  $\nu = \nu^C$

- $\nu_R$  gauge singlets

$\Rightarrow M_R$  not related to SM dynamics, not protected by symmetries

$\Rightarrow M_R \overline{\nu_R^C} \nu_R$  is gauge and Lorentz invariant, renormalisable

- $M_R \overline{\nu_R^C} \nu_R$  violates lepton number conservation  $\Delta L = 2$



# The seesaw mechanism

- $\Delta L = 2, m_\nu \neq 0 \Rightarrow$  New physics
- Seesaw mechanism: New fields with a mass  $M_R >$  EW scale (in general) and Majorana mass terms  
 $\Rightarrow$  Generate  $m_\nu$  in a **renormalizable** way and at tree-level

- Type I seesaw  $\mathcal{L}_{\text{mass}}^{\text{leptons}} = -Y_\ell \bar{L} H \ell_R - Y_\nu \bar{L} \tilde{H} \nu_R - \frac{1}{2} M_R \overline{\nu_R^C} \nu_R + \text{h.c.}$   
 $\Rightarrow$  After EWSB, neutrino mass matrix  $M_{6 \times 6}^\nu$

$$M^\nu = \begin{pmatrix} 0 & m_D \\ m_D^\top & m_R \end{pmatrix} \quad \begin{array}{l} m_D = \nu Y_\nu \text{ Dirac mass matrix} \\ M_R \text{ Majorana mass matrix} \rightarrow \text{Diag}\{m_{R_i}\} \end{array}$$

$\Rightarrow$  Seesaw limit  $M_R \gg m_D$

$$m_\nu^{\text{light}} \approx -m_D M_R^{-1} m_D^\top$$

$$m_\nu^{\text{heavy}} \approx M_R$$

$$\nu^{\text{light}} \approx \nu_L + \nu_L^C$$

$$\nu^{\text{heavy}} \approx \nu_R + \nu_R^C$$

- Charged current matrix  $U_\nu$  is  $3 \times 6$
- 3 charged leptons + 3 light neutrinos:  $\tilde{U}_{\text{PMNS}} = 3 \times 3$  submatrix of  $U_\nu$ , maybe **not unitary**





# Three seesaw mechanisms

- Three minimal tree-level seesaw models  $\Rightarrow$  Three types of heavy fields
  - type I: right-handed neutrinos, SM gauge singlets
  - type II: scalar triplets
  - type III: fermionic triplets

$$m_\nu = -\frac{1}{2} Y_\nu^T \frac{v^2}{M_R} Y_\nu$$

$$m_\nu = -2 Y_\Delta v^2 \frac{\mu_\Delta}{M_\Delta^2}$$

$$m_\nu = -\frac{1}{2} Y_\Sigma^T \frac{v^2}{M_\Sigma} Y_\Sigma$$



# Effective approach to seesaw mechanisms

- Notice that lepton number conservation is **accidental** in the SM (gauge group, field content and renormalizability)
- Majorana  $\nu$  cannot be associated to any global symmetry  
 $\Rightarrow$  Effective non-renormalizable operators
- **Unique** dimension 5 operator for all seesaw mechanisms  
 $\rightarrow$  Violates lepton number  $L \Rightarrow$  **Majorana neutrinos**

$$\delta\mathcal{L}^{d=5} = \frac{1}{2}c_{ij}\frac{(\bar{L}_i\tilde{H})^\dagger(\bar{L}_j\tilde{H})}{\Lambda} + \text{h.c.}$$

- To distinguish the several seesaw mechanisms, either
  - **Directly produce** the heavy states (LHC, LC)
  - Look for **dimension  $\geq 6$  operator effects**  $\rightarrow$  charged lepton flavour violation (cLFV), non-standard interactions, etc



# The inverse seesaw mechanism

- Inverse seesaw  $\Rightarrow$  Consider fermionic gauge singlets  $\nu_{Ri}$  ( $L = +1, i = 1, 2, 3$ ) and  $X_i$  ( $L = +1, i = 1, 2, 3$ )

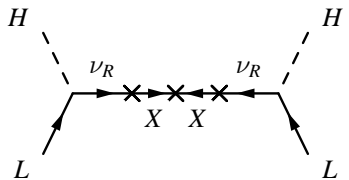
[Mohapatra and Valle, 1986]

$$\mathcal{L}_{inverse} = Y_{\nu}^{ij} \bar{L}_i \tilde{H} \nu_{Rj} - M_R^{ij} \bar{\nu}_{Ri} X_j - \frac{1}{2} \mu_X^{ij} \bar{X}_i^C X_j + \text{h.c.}$$

with  $m_D = Y_{\nu} v, M^{\nu} = \begin{pmatrix} 0 & m_D & 0 \\ m_D^T & 0 & M_R \\ 0 & M_R^T & \mu_X \end{pmatrix}$

$$m_{\nu} \approx \frac{m_D^2 \mu_X}{m_D^2 + M_R^2}$$

$$m_{1,2} \approx \mp \sqrt{m_D^2 + M_R^2} + \frac{M_R^2 \mu_X}{2(m_D^2 + M_R^2)}$$



# The inverse seesaw mechanism

- Type I seesaw:  $M_R \simeq 10^{15}\text{GeV}$  with natural Yukawa  $Y_\nu \sim \mathcal{O}(1)$   
or  $M_R \sim 1\text{TeV}$  with Yukawa  $Y_\nu \sim \mathcal{O}(10^{-6})$   
 $\Rightarrow$  Small active-sterile mixing ( $\frac{m_D}{M_R}$ )  $\rightarrow \tilde{U}_{\text{PMNS}} \sim U_{\text{PMNS}}$   
 $\Rightarrow$  Small deviations from unitarity
- Inverse seesaw:  $M_R \simeq 1\text{TeV}$  with natural Yukawa  $Y_\nu \sim \mathcal{O}(1)$   
 $\Rightarrow$  Large active-sterile mixing  $\rightarrow \tilde{U}_{\text{PMNS}} \neq U_{\text{PMNS}}$   
 $\Rightarrow$  Large deviations from unitarity ?
- Inverse seesaw: **testable at the LHC and low energy experiments**  
Could provide a sterile neutrino at the eV scale (reactor and LSND/MiniBooNE anomalies)



# Lepton flavour universality

- Lepton flavour universality (LFU): gauge boson couplings are **independent of lepton flavours**
- Searches for LFU violation are among the **most precise tests** of the SM

$$\frac{\mathcal{B}(Z^0 \rightarrow \mu^+ \mu^-)}{\mathcal{B}(Z^0 \rightarrow e^+ e^-)} = 1.0009 \pm 0.0028$$

[Schael et al., 2006]

$$\frac{\mathcal{B}(Z^0 \rightarrow \tau^+ \tau^-)}{\mathcal{B}(Z^0 \rightarrow e^+ e^-)} = 1.0019 \pm 0.0032$$

- Deviations from LFU  $\Rightarrow$  **Evidence of New Physics**
- Could help disentangle different neutrino mass generation mechanisms
  - Magnitude of LFU violation
  - Correlation between different LFU tests



# Lepton universality tests

- Couplings to different bosons can be tested:  $\gamma, Z^0, W^\pm$   
Impact of singlet neutrinos  $\Rightarrow$  Focus on  $W^\pm$  couplings
- Many observables can be used
  - Gauge boson decays (e.g.  $W \rightarrow \ell \bar{\nu}$ )
  - Leptonic and semileptonic meson decays (e.g.  $K \rightarrow \ell \bar{\nu}, \bar{B} \rightarrow D \ell^- \bar{\nu}$ )
  - Lepton decays (e.g.  $\ell \rightarrow \ell' \nu \bar{\nu}, \tau \rightarrow K \nu$ )
- We considered light meson decays: pions and kaons  
Decay width plagued by QCD uncertainties  $\Rightarrow$  Consider **ratios**

$$R_P = \frac{\Gamma(P^+ \rightarrow e^+ \nu)}{\Gamma(P^+ \rightarrow \mu^+ \nu)} \quad P = K, \pi$$



# Why $R_K$ and $R_\pi$ ?

- Well measured by the NA62 collaboration [Lazzeroni et al., 2013]:

$$R_K^{\text{exp}} = (2.488 \pm 0.010) \times 10^{-5}$$

Current experimental error:  $\frac{\delta R_K}{R_K} \simeq 0.4\%$

Expected sensitivity:  $\frac{\delta R_K}{R_K} \simeq 0.1\%$

- SM prediction is very precise [Finkemeier, 1996, Cirigliano and Rosell, 2007]:

$$R_K^{\text{SM}} = (2.477 \pm 0.001) \times 10^{-5}$$

- New Physics:  $R_K = R_K^{\text{SM}} (1 + \Delta r_K)$

$$\Delta r_K = (4 \pm 4) \times 10^{-3}$$

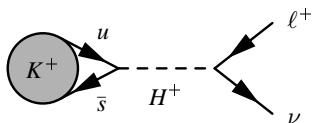
- Similar prospects for  $R_\pi$



# Deviations from the SM

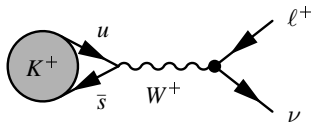
- Origin of LFU violation in  $R_K$ :

- New Lorentz structure in the four-fermion interaction



New fields, new couplings  
e.g. 2 Higgs doublet models,  
Supersymmetry

- Corrections to the SM  $W\ell\nu$  vertex



New states, Higher-order effects  
e.g. Additional neutrinos: low-scale  
seesaw, inverse seesaw





# New Lorentz structure

- Two Higgs Doublet Model (2HDM) [Hou, 1993]

$$\Gamma^{2\text{HDM}}(K^+ \rightarrow \ell^+ \nu) = \Gamma^{\text{SM}}(K^+ \rightarrow \ell^+ \nu) \left( 1 - \tan^2 \beta \frac{m_K^2}{m_{H^+}^2} \frac{m_s}{m_s + m_u} \right)^2$$

$\Rightarrow$  the tree-level correction is **universal**:  $\Delta r_K^{2\text{HDM}} \sim 0$

- Supersymmetry

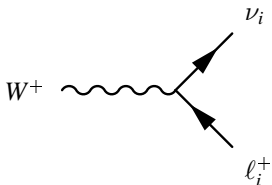
- Tree-level correction is identical to 2HDM  
 $\rightarrow$  higher-order corrections are required [Masiero et al., 2006]
- In the unconstrained MSSM  
 $\mathcal{B}(B \rightarrow \tau \nu)$  limits  $\Delta r_K^{2\text{HDM}} < 10^{-3}$  [Fonseca et al., 2012]



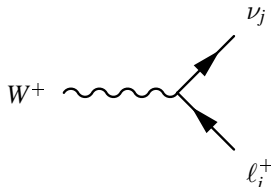
# Modified $W\ell\nu$ vertex

- Naturally arises when leptonic mixing is added to the SM

SM:  $gP_L\delta^{ij}$



SM + massive  $\nu$ :  $gP_L U_\nu^{ij}$



$$i = e, \mu, \tau; \quad j = 1, \dots, n_\nu$$

- If  $n_\nu = 3 \rightarrow U_\nu = U_{\text{PMNS}}$ , unitary
- If  $n_\nu > 3$  (e.g. fermionic singlets)  $\rightarrow U_\nu \neq U_{\text{PMNS}}$ 
  - $\rightarrow U_\nu$  is a  $3 \times n_\nu$  non-unitary matrix
  - $\rightarrow \tilde{U}_{\text{PMNS}}$  is not unitary
- Tree-level corrections to  $R_K$



# Deviation from universality

- Summing over all the kinematically accessible neutrinos (from 1 to  $N_{\max}^{(e)}$ ,  $N_{\max}^{(\mu)}$  the heaviest kinematically allowed neutrino) :

$$R_K = \frac{\sum_{i=1}^{N_{\max}^{(e)}} |U_{\nu}^{1i}|^2 G^{i1}}{\sum_{k=1}^{N_{\max}^{(\mu)}} |U_{\nu}^{2k}|^2 G^{k2}} \quad \text{with}$$

$$G^{ij} = \left[ m_K^2 (m_{\nu_i}^2 + m_{l_j}^2) - (m_{\nu_i}^2 - m_{l_j}^2)^2 \right] \left[ (m_K^2 - m_{l_j}^2 - m_{\nu_i}^2)^2 - 4m_{l_j}^2 m_{\nu_i}^2 \right]^{1/2}$$

- In the SM + 3 massive  $\nu$ , one recovers  $R_K^{SM} = \frac{m_e^2}{m_\mu^2} \frac{(m_K^2 - m_e^2)^2}{(m_K^2 - m_\mu^2)^2}$

- $m_\nu \ll m_\ell \Rightarrow G^{i1} = G^{j1}$

- $U_\nu = U_{\text{PMNS}} \Rightarrow \sum_{i=1}^{n_\nu} |U_{\nu}^{1i}|^2 = (U_\nu U_\nu^\dagger)_{11} = 1$

- 2 ways to deviate from universality:

- (A) sterile neutrinos are lighter than  $m_K$ , with  $m_\nu^{\text{active}} \ll m_{\nu_s} \lesssim m_K$   
 → Phase space effect
- (B) sterile neutrinos are heavier than the kaon,  $m_{\nu_s} > m_K$   
 →  $\tilde{U}_{\text{PMNS}}$  is not unitary



# Is there a visible deviation from LFU ?

- Yes!

JHEP02(2013)048

- Effect of deviation from unitarity [Shrock, 1980, 1981]
- However, current experimental constraints on sterile neutrinos and non-unitary are quite stringent



# Constraints on sterile neutrinos

- Depend on the mass regime and the size of active-sterile mixing
  - Direct searches (e.g. monochromatic lines in  $\pi \rightarrow \mu\nu$ ):  
[Atre et al., 2009, Kusenko, 2009]
  - Non-unitarity constraint  $\tilde{U}_{\text{PMNS}} = (\mathbb{1} - \eta)U_{\text{PMNS}}$ : [Antusch et al., 2009]
  - Lepton flavour violation (e.g.  $\mu \rightarrow e\gamma$ ): [Deppisch and Valle, 2005]
  - $B$  Physics (e.g.  $B \rightarrow \ell\nu$ )



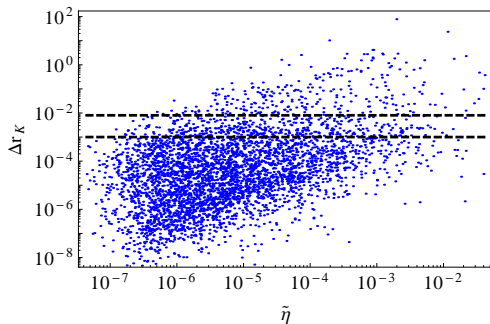
# Constraints on sterile neutrinos

- Depend on the mass regime and the size of active-sterile mixing
  - LHC Higgs searches (e.g. invisible decays):  
[Bhupal Dev et al., 2012, Cely et al., 2013]
  - Electroweak precision data: [del Aguila et al., 2008, Atre et al., 2009]
  - Cosmological observations (e.g. LSS, Lyman- $\alpha$ , CMB, BBN, X-ray): [Smirnov and Zukanovich Funchal, 2006, Kusenko, 2009]  
→ can be evaded with non-standard cosmology (e.g. low reheating temperature [Gelmini et al., 2008])



# $R_K$ in the inverse seesaw

- Inverse seesaw as an illustrative example, only one among other possibilities
- Numerical results in scenario (A):  $m_\nu^{\text{active}} \ll m_{\nu_s} \lesssim m_K$



$$M_R \in [0.1 \text{ MeV}, 200 \text{ MeV}]$$

$$\mu_X \in [0.01 \text{ eV}, 1 \text{ MeV}]$$

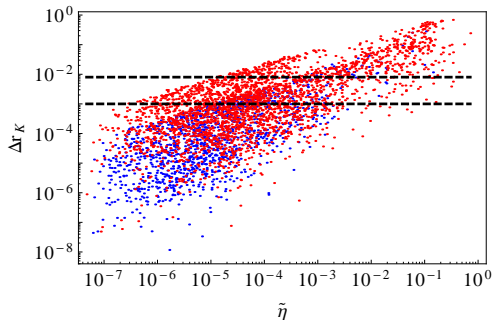
$$\tilde{\eta} = 1 - |\text{Det}(\tilde{U}_{\text{PMNS}})|$$

- Comply with all constraints, except cosmological bounds
- Large LFU violation  $\Delta r_K \sim 1$  can be reached
- Small  $Y_\nu \rightarrow$  no cLFV signal expected



# $R_K$ in the inverse seesaw

- Numerical results in scenario (B):  $m_{\nu_s} > m_K$



$$M_R \in [1 \text{ GeV}, 10^6 \text{ GeV}]$$

$$\mu_X \in [0.01 \text{ eV}, 1 \text{ MeV}]$$

$$Y_\nu > 10^{-2} \quad Y_\nu < 10^{-2}$$

$$\tilde{\eta} = 1 - |\text{Det}(\tilde{U}_{\text{PMNS}})|$$

- Comply with all constraints, even the stringent non-unitarity bounds
- Large LFU violation  $\Delta r_K \sim 1$  can be reached
- Large  $Y_\nu \rightarrow \mathcal{B}(\mu \rightarrow e\gamma)$  is within MEG reach
- Specific to the inverse seesaw with its large active-sterile mixing





# $R_K$ summary

- Source: **modified  $W\ell\nu$  vertex** from extra sterile neutrinos
- Mechanism: **phase space effect** for  $m_\nu^{\text{active}} \ll m_{\nu_s} \lesssim m_K$   
**non-unitarity** of  $\tilde{U}_{\text{PMNS}}$  for  $m_{\nu_s} > m_K$
- Results: large LFU violation ( $\Delta r_K \lesssim 1$ ) in the inverse seesaw
- **Similar results** for  $R_\pi$



# Other observables

- Leptonic  $W^\pm$  decays

$$R_{\tau\ell}^W = \frac{2\mathcal{B}(W \rightarrow \tau\bar{\nu}_\tau)}{\mathcal{B}(W \rightarrow \mu\bar{\nu}_\mu) + \mathcal{B}(W \rightarrow e\bar{\nu}_e)} = 1.077 \pm 0.026$$

$\Rightarrow 2.8\sigma$  deviation from the SM prediction  $R_{\tau\ell}^W|_{SM} = 0.999$   
[LEP, 2005, Kniehl et al., 2000]

- Other leptonic meson decays

$$R_{D_s} = \frac{\Gamma(D_s^+ \rightarrow \tau^+\nu)}{\Gamma(D_s^+ \rightarrow \mu^+\nu)} \simeq 9.2$$

$\Rightarrow$  roughly  $1\sigma$  away from the SM prediction  $R_{D_s}|_{SM} \simeq 10.1$   
[Beringer et al., 2012, Charles et al., 2011]



# Other observables

- Semileptonic  $\tau$  decays

$$R_{K,\mu}^\tau = \frac{\mathcal{B}(\tau^- \rightarrow K^- \nu)}{\mathcal{B}(K^+ \rightarrow \mu^+ \nu)} \simeq 469$$

$\Rightarrow$  within  $1\sigma$  of the SM prediction  $R_{K,\mu}|_{SM} \simeq 474$  [Beringer et al., 2012]

- Semileptonic meson decays

$$R(D) = \frac{\mathcal{B}(\bar{B} \rightarrow D \tau^- \bar{\nu}_\tau)}{\mathcal{B}(\bar{B} \rightarrow D \ell^- \bar{\nu}_\ell)} = 0.440 \pm 0.072$$

$\Rightarrow$  deviates by  $1.7\sigma$  from the SM prediction  $R(D)|_{SM} = 0.31 \pm 0.02$   
[Lees et al., 2012, Becirevic et al., 2012]



# Other observables

- 3-body lepton decays:  $\ell_i \rightarrow \ell_j \nu \nu$

- Test LFU via ratios, e.g.

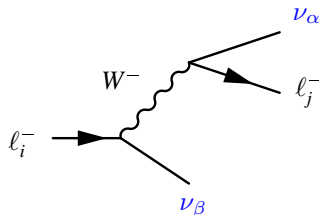
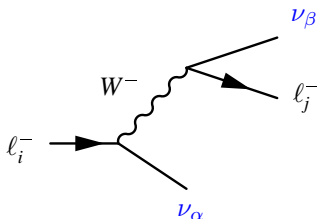
$$R_\tau = \frac{\Gamma(\tau^- \rightarrow \mu^- \nu \nu)}{\Gamma(\tau^- \rightarrow e^- \nu \nu)} = 0.9764 \pm 0.0030$$

$\Rightarrow$  within  $2\sigma$  of the SM prediction  $R_\tau|_{SM} \simeq 0.9726$  [Beringer et al., 2012]

- 2 massive neutrinos in the final state  
 $\rightarrow$  Sensitive to the Majorana or Dirac nature of the neutrino



# Other observables



- $\Gamma_{\text{Majorana}} = \Gamma_{\text{Dirac}} + \Gamma_{\text{interferences}}$

$$\Gamma_{\text{interferences}} \propto \Re(U_{i\alpha}^* U_{j\beta} U_{i\beta} U_{j\alpha}^*) \frac{m_{\nu_\alpha} m_{\nu_\beta}}{m_{\ell_i}^2}$$

- Large active-sterile mixing + heavy sterile neutrinos ( $m_{\nu_s} \sim \mathcal{O}(m_{\ell_i})$ )  
 $\Rightarrow$  Potentially sizeable deviations (under investigation)



# Conclusion

- Lepton universality tests: **good theoretical** precision and **experimental prospects**
- Sterile neutrinos can lead to a large violation of LFU at **tree-level**
- $R_K$  particularly well-suited for this search
- **Large deviations** from the SM can be found
- Can appear in other **observables with leptonic charged currents**  
→ Currently under investigation



Backup slides



# Leptogenesis

- Generate the baryonic (leptonic) asymmetry  $\rightarrow$  Sakharov conditions [Sakharov, 1967]
  - Out of equilibrium process
  - Baryon (lepton) number violation
  - C and CP violation
- Impossible in the Standard Model: not enough CP violation [Gavela et al., 1994]
- Use the leptonic sector
  - Majorana mass term violates lepton number conservation  $\Rightarrow$  Passed to the baryonic sector via sphalerons ( $B - L$  conserving)
  - Neutrinos mass matrix  $\Rightarrow$  Extra sources of CP violation ( $\delta_{13}, \alpha_{1,2}$ )





# Scan method

- Inverse seesaw as an illustrative example, only one among other possibilities
- Random scan on  $M_R$  and  $\mu_X$  entries
- $Y_\nu$  obtained from neutrino data via the Casas-Ibarra parametrization [Casas and Ibarra, 2001]

$$Y_\nu^T = \frac{\sqrt{2}}{v} V^\dagger \text{diag}(\sqrt{m_1}, \sqrt{m_2}, \sqrt{m_3}) R \text{diag}(\sqrt{M_1}, \sqrt{M_2}, \sqrt{M_3}) U_{PMNS}^\dagger$$

where  $R$  is a complex orthogonal matrix and  $V$  a unitary matrix that decompose  $M = M_R \mu_X^{-1} M_R^T$  according to  $M = V^\dagger \text{diag}(M_1, M_2, M_3) V^*$ .

- Apply the constrains



